Homogeneous Finite-Time Pose Tracking of Leader–Following Spacecraft Formation Using a Twistor-Based Model

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Abstract: For integrated pose control of six-degree-of-freedom (6-DOF) leader–following spacecraft formations, the coupled relative pose between two spacecraft is described by a unified model in the framework of twistors, based on which a finite-time control law is devised for the 6-DOF dynamic system. Firstly, some necessary coordinate frames are defined, and then, relative 6-DOF dynamics of the follower spacecraft with respect to the desired frame are modeled by using twistors. Secondly, an integrated pose controller is designed for the 6-DOF formation that takes full advantage of homogeneous theory to guarantee finite-time stability. Finally, numerical simulations are carried out to validate the effectiveness of the proposed controller. Developing integrated 6-DOF formation control law based on twistors is more straightforward than conventional methods, and the finite-time algorithm achieves stronger robustness than asymptotic ones.

Keywords: 6-DOF spacecraft formation; twistor; homogeneous theory; finite-time control; integrated pose control

1. Introduction

Due to the increasingly demanding requirements of future space missions, space technology is encountering an increasing number of challenges as it progresses [1]. More and more attention is being paid to 6-DOF spacecraft formations because plenty of low-cost satellites are being sent to space to work cooperatively [2]. Ref. [3] researched the coordinated control of 6-DOF spacecraft formations with external disturbances and communication delays considered in the situation for which velocity is immeasurable. In Ref. [4], a coupled pose error dynamic model of 6-DOF spacecraft formations was established, based on which the authors devised a pose tracking law by employing the sliding mode control theory with an event-triggering strategy introduced to reduce the communication rate between the spacecraft. In Ref. [5], a fixed-time relative pose tracking controller with hysteretic quantized control input was devised for 6-DOF spacecraft formations, in which the fast terminal sliding mode control method was used in conjunction with an adaptive law. The proposed control scheme guaranteed fixed-time convergence of the closed-loop system in the presence of input quantization, uncertainties, and disturbances. Meanwhile, it lightened the communication burden and avoided the oscillations resulting from the logarithm quantizer.

In conventional methods for modeling spacecraft dynamics, translational motion and rotational motion are represented separately [6]. However, ignoring the coupling between translation and rotation may lead to nonnegligible errors in spacecraft formations [7]. Dual quaternions, which originate from Clifford algebra, can describe the translation and rotation...
of a rigid body simultaneously and briefly [8,9]. For the leader–following formation of two spacecraft, the relative pose of the spacecraft can be represented in a unified way by employing dual quaternions, which can simplify the design of the control law for 6-DOF formations [10]. Li et al. [11] investigated the coupled pose control of 6-DOF spacecraft motion using dual quaternions, and an event-triggered PI-like control algorithm with an active disturbance rejection strategy was proposed for pose tracking, which considers the separate time scale of the 6-DOF system and thus possesses a lower communication rate than that of controllers using a coupled-triggering strategy. A dual-quaternion-based extended Kalman filter was proposed first in Ref. [12] to estimate both the pose and (linear and angular) velocity of a spacecraft.

Another tool for uniformly describing the coupled pose motion is twistors, which were originally derived by applying the principles of geometric algebra [13,14]. Although dual quaternions can describe the translational and rotational motion of spacecraft in a unified way, dual-quaternion-based representation possesses redundant elements that may result in additional communication and computation burden. Further, a dual quaternion used to represent pose motion must be a unit quaternion, which is inconvenient in some aspects [13]. For example, it is difficult to apply an unscented Kalman filter to a dual-quaternion-represented system [15]. In Ref. [14], twistors, which are of six dimensions, were utilized to parameterize translational and rotational motion of elastically coupled rigid bodies. The definition of twistors has been given in the framework of geometric algebra as well. Ref. [13] presented a coupled relative pose dynamic model of spacecraft based on twistors and then designed an asymptotically stable PD-like controller. Zhang and Cai [16] used twistors to describe the pose motion of a spacecraft relative to an asteroid, and a collision-free pose control scheme for landing was proposed with the help of the artificial potential field (APF) technique. Ref. [17] further developed an adaptive 6-DOF saturated control algorithm for spacecraft landing on an asteroid on the basis of a twistor model. Similarly, the APF technique was introduced into the backstepping method to avoid collisions. Zhang and Li [18] devised a twistor-based control scheme for spacecraft formations considering the coupling between position and attitude dynamics and achieved finite-time stability by the application of the fast terminal sliding mode technique. In Ref. [19], the 6-DOF distributed control of a spacecraft formation near an asteroid was investigated in the framework of twistors to cope with position–attitude coupling. To overcome the numerical difficulty, scaled twistors were introduced into the relative dynamic model, and then the consensus theory was combined with the dynamic surface control method to derive a distributed control protocol.

Due to its fast response and strong robustness [20,21], finite-time control in the field of spacecraft control is becoming a hot research topic [22,23]. In Ref. [24], a non-linear quaternion-based finite-time control scheme with bounded control inputs was developed with the aid of homogeneous theory, and the unwinding phenomenon was avoided by selecting appropriate equilibria. Gui et al. [22] constructed a finite-time output feedback pose tracking control scheme for spacecraft based on dual quaternion representation, where a simple second-order filter was designed to provide the estimates of spacecraft velocity. In the control scheme, the homogeneous theory was employed to ensure finite-time stability, and the hybrid system theory was used to achieve global convergence. Ref. [25] presented a finite-time pose tracking algorithm without velocity measurements in the framework of dual quaternions. Likewise, the unavailable velocity was estimated by an observer, but the finite-time stability of the closed-loop system was guaranteed via the terminal sliding mode control theory. To avoid collisions in spacecraft formations, Zhang et al. [26] designed a finite-time controller for satellite formations by embedding an APF into the second-order sliding mode control method.

Though finite-time control of spacecraft has been extensively studied, twistor-based integrated position and attitude control, which possesses the advantages of both twistor representation and finite-time convergence, has not appeared (the approach in Ref. [18] is just practical finite-time stable). Therefore, investigating finite-time control of spacecraft
formations in the framework of twistors will greatly promote the development of spacecraft formations. The existing challenges in representative works are summarized in Table 1. For integrated position and attitude control of 6-DOF spacecraft formations, a finite-time control law based on twistor representation is proposed to get rid of the normalization constraint and state dimension redundancy in order to improve the response speed and robustness.

Table 1. Challenges in representative works on spacecraft control.

<table>
<thead>
<tr>
<th>Control Law</th>
<th>State Representation</th>
<th>Main Challenge</th>
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</thead>
<tbody>
<tr>
<td>Control law in [27]</td>
<td>quaternions</td>
<td>only attitude control</td>
</tr>
<tr>
<td>Control law in [22]</td>
<td>dual quaternions</td>
<td>state redundancy and normalization constraint</td>
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<td>Control law in [19]</td>
<td>twistors</td>
<td>asymptotic stability</td>
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<tr>
<td>Control law in [18]</td>
<td>twistors</td>
<td>practical finite-time stability</td>
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This paper presents a finite-time pose control law for leader–following spacecraft formations with the application of the homogeneous theory based on the coupled relative pose dynamic model described by twistors. First, the relative pose dynamics of spacecraft formations with the application of the homogeneous theory and Lyapunov stability theory. Next, case simulations are carried out to confirm the effectiveness of the devised control algorithm with necessary discussions and analysis presented. The advantages of the control law are clearly shown through comparisons. Finally, some conclusions are given to finalize this paper.

2. Preliminaries

Before proceeding to the main results, this section presents some critical mathematical preliminaries to ease the understanding of the proposed algorithm.

2.1. Definitions and Lemmas

Let $\text{sign}^n(y) = [|y_1|^\alpha \text{sign}(y_1), |y_2|^\alpha \text{sign}(y_2), \ldots, |y_n|^\alpha \text{sign}(y_n)]^T$ for any $y = [y_1, y_2, \ldots, y_n]^T \in \mathbb{R}^n$ and $\alpha \geq 0$, where $\text{sign} (\bullet)$ is the sign function. Clearly, the function $\text{sign}^n(y)$ is non-smooth and continuous if $\alpha \in (0,1)$. Let $\mu > 0$ and a vector $h = [h_1, h_2, \ldots, h_n] \in \mathbb{R}^n$, where $h_i > 0$, $i = 1, 2, \ldots, n$. For any $y \in \mathbb{R}^n$, define an operator $\Lambda^h_\mu$ as $\Lambda^h_\mu y = [\mu^{h_1}y_1, \mu^{h_2}y_2, \ldots, \mu^{h_n}y_n]^T$.

**Definition 1** ([28]). Assume the following system

$$ y = g(y, t), \ g(0, t) = 0, \ x \in \mathbb{R}^n, \quad (1) $$

where $g(y, t) = [g_1(y, t), \ldots, g_n(y, t)]^T$ is continuous on an open neighborhood of the origin. For $i = 1, 2, \ldots, n$ and any $\mu > 0$, if $g_i(\Lambda^h_\mu y, t) = \mu^{h_i+k}g_i(y, t)$, $g(y, t)$ is called a homogeneous function of degree $k \in \mathbb{R}$ with regard to $\Lambda^h_\mu$. Then, the system (1) is homogeneous if $g(y, t)$ is homogeneous.

**Definition 2** ([29]). Let $\Xi$ denote a neighborhood of $y = 0$. If the system (1) is uniformly Lyapunov stable on $\Xi$ and converges uniformly to the origin in finite time on $\Xi$, the origin is uniformly finite-time stable (UFTS).

**Lemma 1** ([30]). Assume the system represented by

$$ \dot{y} = g(y) + \bar{g}(y, t), \ g(0) = 0, \ y \in \mathbb{R}^n, \quad (2) $$
where $g(y)$ is a homogeneous continuous function of degree $k < 0$ with regard to $A^k_i$ and $g(y, t)$ meets the condition $g(0, t) = 0$. It is assumed that the system $\dot{y} = g(y)$ asymptotically converges to the equilibrium $y = 0$. Then, the system (2) is uniformly locally finite-time stable with $y = 0$ being its equilibrium if $g(y, t)$ satisfies

$$\lim_{\mu \to 0} \frac{\delta_i(A^k_i y, t)}{\mu^{n+k}} = 0, \ i = 1, 2, \ldots, n, \ \forall y \neq 0. \quad (3)$$

The proof of Lemma 1 can be seen in [27]. As shown in [27], if the system (2) is disturbed by a small disturbance $g_d(t)$ and $\dot{y} = g(y) + g_d(y, t) + g_d(t)$, $y$ is still finite-time stable.

2.2. Coordinate Frame Definition and Twistor-Based Motion Equations

To facilitate describing the 6-DOF dynamics of spacecraft, some coordinate frames are defined as follows. The right-handed inertial coordinate frame is fixed in the inertial space and denoted by $O_X X_Y Y_Z (F_I)$. The origin of the coordinate frame $O_X X_Y Y_Z (F_B)$ is the barycenter of the follower spacecraft, with the axes aligned with the principal axes of inertia to constitute a right-handed coordinate system. The desired pose of the follower spacecraft is described by the coordinate frame $O_X X_Y Y_Z (F_D)$, which means that $F_D$ is the desired frame of $F_B$. The right-handed frame $O_Y Y_T Y_Z (F_T)$ is fixed with the leader spacecraft, the origin of which is located at the barycenter and the axes of which are aligned with its principal inertia axes. The relationship between these coordinate frames is illustrated in Figure 1.

![Figure 1. Relationship between the frames $F_I$, $F_T$, $F_D$, and $F_B$.](image)

It is expected that the axes of $F_B$ coincide with the axes of $F_D$. The pose of $F_B$ relative to $F_D$ described in the framework of dual quaternions is given by [31]

$$\tilde{q}_{BD} = q_{BD} + \frac{1}{2} \epsilon r^D_{BD} q_{BD} = q_{BD} + \frac{1}{2} \epsilon q_{BD} r^B_{BD}, \quad q_{BD} \in \mathbb{Q}, \ r^D_{BD}, r^B_{BD} \in \mathbb{Q}^p \quad (4)$$

where $\mathbb{Q}$ and $\mathbb{Q}^p$ denote the set of quaternions and the set of quaternions with null scalar parts, respectively, $q_{BD}$ is the quaternion that represents the attitude of $F_B$ relative to $F_D$, $r^D_{BD}$ is the displacement vector of $O_B$ (the origin of $F_B$) relative to $O_D$ (the origin of $F_D$) resolved in the coordinate system $F_D$, and $r^B_{BD}$ is the same vector resolved in $F_B$. The term $\epsilon$, satisfying $\epsilon \neq 0$ and $\epsilon^2 = 0$, is the dual unit. Thus, the twistor, which represents the pose of $F_B$ relative to $F_D$, is expressed by the dual quaternion $\tilde{q}_{BD}$ as [13,16]

$$\tilde{B}_{BD} = \tilde{q}_{BD} - \frac{1}{\tilde{q}_{BD} + 1} = p_{BD} + \epsilon b_{BD}, \quad \tilde{q}_{BD}, \tilde{B}_{BD} \in \mathbb{D}, \ p_{BD}, b_{BD} \in \mathbb{Q}^p \quad (5)$$

where $\mathbb{D}$ is the set of dual quaternions, $\tilde{1}$ is the unit dual quaternion, $p_{BD}$ (the real part of $\tilde{B}_{BD}$) are the Modified Rodrigues Parameters (MRPs) of the body-fixed frame $F_B$ relative to its desired frame $F_D$, and the term $b_{BD}$ is the dual part relying on the MRPs and the relative position vector. Hence, a twistor can also be regarded as dual MRPs [13].
\( \mathbf{p}_{BD} \) and \( \mathbf{b}_{BD} \) are the special quaternions with zero scalar parts, and the expressions of their vector parts are given by

\[
\mathbf{p}_{BD} = \frac{\mathbf{q}_v}{1 + \mathbf{q}_s},
\]

(6)

\[
\mathbf{b}_{BD} = \mathbf{f}_{BD} = \Gamma \mathbf{r}_{BD}^B = \Gamma \mathbf{r}_{BD}^B,
\]

(7)

where

\[
\mathbf{q}_v = ((\mathbf{q}_{BD})_{real})_{vec}, \quad \mathbf{q}_s = ((\mathbf{q}_{BD})_{real})_{scalar}
\]

(8)

\[
\Gamma = \frac{1}{4} \left( 1 - \frac{1}{2} \mathbf{p}_{BD} \mathbf{p}_{BD}^T \right) I_{3 \times 3} - \frac{1}{2} \mathbf{p}_{BD}^T + \frac{1}{2} \mathbf{p}_{BD} \mathbf{p}_{BD}^T,
\]

(9)

where (\( \bullet \)) symbolizes the vector part of a quaternion; (\( \bullet \)\_real, (\( \bullet \)\_vec, and (\( \bullet \)\_scalar) extract the real part of a dual quaternion, vector part of a quaternion, and scalar part of a quaternion, respectively; and (\( \bullet \)^\* gives a 3 \( \times \) 3 skew-symmetric matrix that satisfies \( \mathbf{a} \times \mathbf{b} = (\mathbf{a})^\* \mathbf{b} \), \( \mathbf{a}, \mathbf{b} \in \mathbb{R}^3 \). The above equations show that a twistor comprises two quaternions whose scalar parts are zero, indicating that the pose is described by six parameters. Clearly, the dimension of twistors is less than that of dual quaternions. More information about twistors can be obtained from Refs. [13,14].

The relative pose motion is described by twistors as [13]

\[
\hat{\mathbf{b}}_{BD} = \frac{1}{4} \left( \hat{\mathbf{1}} + \mathbf{B}_{BD} \right) \mathbf{V}_{BD}^B \left( \hat{\mathbf{1}} - \mathbf{B}_{BD} \right),
\]

(10)

\[
\mathbf{V}_{BD}^B = \mathbf{M}^{-1} \left[ \hat{\mathbf{F}}_{BD}^B + \hat{\mathbf{F}}_{BD}^B \times \left( \hat{\mathbf{M}} \hat{\mathbf{V}}_{BD}^B \right) \right] - \frac{1 - \mathbf{B}_{BD}}{1 + \mathbf{B}_{BD}} \hat{\mathbf{V}}_{BD}^B \frac{\mathbf{1} + \mathbf{B}_{BD}}{1 - \mathbf{B}_{BD}} - \frac{1}{1 + \mathbf{B}_{BD}} \hat{\mathbf{V}}_{BD}^B \frac{\mathbf{1} + \mathbf{B}_{BD}}{1 - \mathbf{B}_{BD}}
\]

(11)

where \( \mathbf{V}_{BD}^B = \omega_{BD}^B + \epsilon (\mathbf{r}_{BD}^B + \omega_{BD}^B \times \mathbf{r}_{BD}^B) \), with \( \omega_{BD}^B \) and \( \mathbf{r}_{BD}^B \) being the angular velocity vector and position vector, respectively, of \( \mathbf{F} \) with respect to \( \mathbf{F} \) resolved in \( \mathbf{F} \). The expressions of \( \mathbf{V}_{BD}^D \) and \( \mathbf{V}_{BD}^F \) resemble that of \( \mathbf{V}_{BD}^B \) and are omitted here for conciseness. According to the definition of twistors, \( \mathbf{B}_{BD} \) and \( (\mathbf{B}_{BD})^{-1} \) represent the same relative pose between frames. As shown in Ref. [13], when \( \mathbf{q}_{BD} = -\mathbf{1}, \mathbf{B}_{BD} \) is singular and \( (\mathbf{B}_{BD})^{-1} \) is used instead of \( \mathbf{B}_{BD} \). The term \( \mathbf{M} \) denotes the dual inertia matrix, which is given by [16]

\[
\mathbf{M} = m \frac{d}{d\epsilon} I_{3 \times 3} + \epsilon \mathbf{J},
\]

(12)

where \( \mathbf{J} \) is the inertia matrix of the follower spacecraft, and \( m \) is its mass. Then, the inversion of \( \mathbf{M} \) is

\[
\mathbf{M}^{-1} = \frac{d}{d\epsilon} \mathbf{J}^{-1} + \epsilon \frac{1}{m} I_{3 \times 3}.
\]

(13)

The dual force caused by gravity expressed in \( \mathbf{F} \) is defined as \( \hat{\mathbf{F}}_s^B = \mathbf{f}_s^B + \epsilon \mathbf{r}_s^B \). The terms \( \mathbf{f}_s^B \) and \( \mathbf{r}_s^B \) represent the gravitational force vector and gravitational gradient torque vector resolved in \( \mathbf{F} \). The term \( \hat{\mathbf{f}}_u^B \) is the dual control force expressed as \( \hat{\mathbf{f}}_u^B = \mathbf{f}_u^B + \epsilon \mathbf{r}_u^B \), where \( \mathbf{r}_u^B \) and \( \mathbf{f}_u^B \) are the control torque and force, respectively, resolved in \( \mathbf{F} \).
3. Twistor-Based Finite-Time Control Law Design

For the convenience of twistor-based finite-time control law design using homogeneous theory, two functions of \( \hat{B}_{BD} \) are given by

\[
\hat{p}_1(\hat{B}_{BD}, 1 - \alpha_1) = \begin{cases} 
\frac{(1 + \hat{B}_{BD}^\wedge [1 - \hat{B}_{BD}])^T}{4(\hat{B}_{BD} \circ \hat{B}_{BD})^{1 - \alpha_1}} & \in \mathbb{T}, \quad \Vert \hat{B}_{BD} \Vert \neq 0, \\
\hat{\theta} & \in \mathbb{T}, \quad \Vert \hat{B}_{BD} \Vert = 0
\end{cases}
\]

\[
\hat{p}_2(\hat{B}_{BD}, 1 - \alpha_1) = \begin{cases} 
\frac{\hat{B}_{BD}}{4(\hat{B}_{BD} \circ \hat{B}_{BD})^{1 - \alpha_1}} & \in \mathbb{T}, \quad \Vert \hat{B}_{BD} \Vert \neq 0, \\
\hat{\theta} & \in \mathbb{T}, \quad \Vert \hat{B}_{BD} \Vert = 0
\end{cases}
\]

where \( \mathbb{T} \) is the set of twistors; \( \hat{\theta} \) denotes a twistor with all zero elements; \( \circ \) represents circle multiplication between twistors; \( \bullet^\wedge \) and \( \bullet^\vee \) are the matrices attained from multiplication of twistors, which satisfy \( \hat{X}^\wedge \hat{Y} = [\hat{X}]^\wedge [\hat{Y}]^\vee, \hat{X}, \hat{Y} \in \mathbb{T} \) [16,17,19]; the operation \( \Vert \bullet \Vert \) denotes the Euclidean norm; and \( 0 \leq \alpha < 1 \). It can be proved that the two functions in Equations (14) and (15) are continuous. For \( 0 < \alpha_1 < 1, \alpha_2 > 0, k_1 > 0, \) and \( k_2 > 0, \) a twistor-based control scheme with finite-time stability is designated as

\[
\dot{F}^B_n = -k_1\hat{p}_1(\hat{B}_{BD}, 1 - \alpha_1) + k_2\text{sign}^\alpha(\hat{V}_{BD}^B) \dot{F}^B_s + \hat{V}_{BI}^B \times (\hat{M}\hat{V}_{BI}^B)
\]

\[
= \frac{1}{2}\hat{V}^B \circ (\hat{M}\hat{V}^B) + \frac{2k_1}{1 + \alpha_1} \kappa(\hat{B}, 1 + \alpha_1)
\]

where \( \kappa(\hat{B}, 1 + \alpha_1) = (\hat{B} \circ \hat{B})^{(1 + \alpha_1)/2} \). \( V(x) \) satisfies \( V(x) \geq 0 \) and \( V(x) = 0 \) if and only if \( x \in O \). The time derivative of \( V(x) \) is given by

\[
\dot{V}(x) = \hat{V}^B \circ (\hat{M}\hat{V}^B) + \frac{2k_1}{1 + \alpha_1} \kappa(\hat{B}, 1 + \alpha_1)
\]

\[
= \hat{V}^B \circ [k_1\hat{p}_1(\hat{B}, 1 - \alpha_1) + k_2\text{sign}^\alpha(\hat{V}^B)]
\]

\[
= -\hat{V}^B \circ \frac{k_2\text{sign}^\alpha(\hat{V}^B)}{4(\hat{B} \circ \hat{B})^{1 - \alpha_1}} < 0, \forall \hat{V}^B \neq \hat{\theta}.
\]

Therefore, \( O \) is uniformly Lyapunov stable.

Equation (18) implies that \( \lim_{t \to \infty} V(t) = V(\infty) \) exists and is finite, and therefore \( \hat{B} \hat{V}^B \) are uniformly bounded. Because \( \hat{V}^B_B \) and \( \hat{V}^B_B \) are set to be bounded, \( \hat{V}^B \) is uniformly bounded according to Equations (11) and (16). Thus, \( V(x) \) is uniformly continuous as a result. Then, the conclusion that \( \hat{V}(x) \to 0 \) and thus \( \hat{V}^B \to 0 \) is obtained through Barbalat’s
lemma [32]. Since both $\hat{B}$ and $\hat{V}^B$ are uniformly bounded and uniformly continuous and $\hat{\rho}_1(\hat{B}, 1 - \alpha_1)$ is a continuous function of $\hat{B}$, then $\hat{\rho}_1(\hat{B}, 1 - \alpha_1)$ is uniformly continuous with respect to time. Substituting Equation (16) into (11), it follows that $\dot{\Lambda}$ can be obtained readily by some simple operations:

$$\dot{\Lambda} = \frac{1}{4} \hat{V}^B$$

where

$$f_1 = \frac{1}{4} \hat{V}^B$$

$$f_2 = -\left[ k_1 \hat{\rho}_2 \left( \hat{B}, 1 - \alpha_1 \right) + k_2 \text{sign} \left( \hat{V}^B \right) \right]^s,$$

$$f_1(x, t) = \frac{1}{4} \left( 1 + \hat{B} \right) \hat{V}^B \left( 1 - \hat{B} \right) - \frac{1}{4} \hat{V}^B,$$

$$f_2(x, t) = k_1 \left[ \hat{\rho}_2 \left( \hat{B}, 1 - \alpha_1 \right) - \hat{\rho}_1 \left( \hat{B}, 1 - \alpha_1 \right) \right]^s.$$}

Note that Equations (19) and (20) take the same form as the system in (2). With the same Lyapunov function candidate as Equation (17), it can be proven by similar analysis that the set $O$ possesses asymptotic stability for the system composed of Equations (19) and (20). Let $\Lambda^B_x = (\mu h^1 \hat{B}, \mu h^2 \hat{V}^B)$ with the dilation $\Lambda^B_x$. For any $k < 0$, $\alpha_2 = 2\alpha_1/(1 + \alpha_1)$ and $\alpha_1 \in (0, 1)$, set $b_1 = -2k/(1 - \alpha_1)$, $b_2 = -k(1 + \alpha_1)/(1 - \alpha_1)$. Then, the following equations can be obtained readily by some simple operations:

$$f_1 \left( \Lambda^B_x, t \right) = \frac{1}{4} \mu h^2 \hat{V}^B = \frac{1}{4} \mu h^1 + k \hat{V}^B$$

$$f_2 \left( \Lambda^B_x, t \right) = -\left[ k_1 \hat{\rho}_2 \left( \mu h^1 \hat{B}, 1 - \alpha_1 \right) + k_2 \text{sign} \left( \mu h^2 \hat{V}^B \right) \right]^s$$

$$= -\mu h^1 + k \left[ k_1 \hat{\rho}_2 \left( \hat{B}, 1 - \alpha_1 \right) + k_2 \text{sign} \left( \hat{V}^B \right) \right]^s.$$

According to Definition 1, the vector field $f = [f_1, f_2]^T$ has homogeneity of degree $k$ with regard to the dilation $\Lambda^B_x$. For $f_1(x, t)$ and $f_2(x, t)$, we have

$$\lim_{\mu \to 0} \frac{f_1 \left( \Lambda^B_x, t \right)}{\mu h^1 + k} = \lim_{\mu \to 0} \left[ \left( 1 + \mu h^1 \hat{B} \right) \left[ 1 - \mu h^1 \hat{B} \right]^v - I_{8 \times 8} \right] \mu h^2 \hat{V}^B = 0,$$

$$\lim_{\mu \to 0} \frac{f_2 \left( \Lambda^B_x, t \right)}{\mu h^2 + k} = \lim_{\mu \to 0} \left[ k_1 \left[ \hat{\rho}_2 \left( \mu h^1 \hat{B}, 1 - \alpha_1 \right) - \hat{\rho}_1 \left( \mu h^1 \hat{B}, 1 - \alpha_1 \right) \right]^s \right]$$

$$= \lim_{\mu \to 0} \frac{k_1 \mu h^1 \left[ I_{8 \times 8} - \left( \left[ 1 + \mu h^1 \hat{B} \right] \left[ 1 - \mu h^1 \hat{B} \right]^T \right] \hat{B} \right]^s}{4 \mu h^2 \mu h^1 \hat{B}^{1 - \alpha_1}} = 0.$$
One can confirm the uniformly local finite-time stability of the equilibrium by applying Lemma 1 to Equations (25)–(28). Therefore, the compact set $O$ is UFTS by summarizing the above analysis.

4. Simulations

The proposed method is simulated and analyzed in this section to verify its effectiveness. Further, the simulation results of the PD-like method in [13] are also shown for comparison. For the simulations of the above two methods, the total simulation time is 200 s and the time step is set to 0.01 s. Figure 2 describes how the simulation model is set up. The inertia matrix of the spacecraft is set to

$$J = \begin{bmatrix} 22.0 & 1.0 & 0.5 \\ 1.0 & 20.0 & 0.8 \\ 0.5 & 0.8 & 25.0 \end{bmatrix} \text{kg} \cdot \text{m}^2,$$  

with its mass being $m = 50$ kg.

The desired position vector of $F_D$ relative to $F_T$ expressed in $F_T$ is $\vec{x}_{DT}^T = [12, 0, 0]^T$ m. The scenario is to keep $F_D$ stationary relative to $F_T$, which means $\vec{v}_{DT}^T = [0, 0, 0]^T$ m/s. Quaternion $q_{DT} = t_i + t_j + 0.5k - 0.8660$ represents the desired attitude of $F_D$ relative to $F_T$. The desired angular velocity of $F_D$ with respect to $F_T$ resolved in $F_T$ is $\vec{\omega}_{DT}^T = [0, 0, 0]^T$ rad/s. The initial position vector and velocity vector of $F_B$ relative to $F_D$ expressed in $F_B$ are $\vec{x}_{BD}^B = [-2, 5, 3]^T$ m and $\vec{v}_{BD}^B = [0.1, 0.05, -0.2]^T$ m/s, respectively. The initial relative angular velocity of $F_B$ relative to $F_D$ expressed in $F_B$ is $\vec{\omega}_{BD}^B = [0.87, 0.17, 1.22]^{T} \times 10^{-2}$ rad/s, and the initial quaternion of $F_B$ relative to $F_D$ is $q_{BD} = 0.1917t + 0.7999j + 0.3320k + 0.4618$. Table 2 displays the orbital parameters of the leader spacecraft. Assume that $F_T$ always coincides with the local vertical and local horizontal coordinate frame. External disturbances are given by

$$\vec{f}_{D}^B = \left(\begin{array}{c} 2 \sin(0.1t) + 0.2 \sin(10t + \frac{\pi}{4}) \\ 5 \sin(0.1t + \frac{\pi}{4}) + 0.5 \sin(10t + \frac{\pi}{4}) \\ 4 \sin(0.1t + \frac{\pi}{4}) + 0.4 \sin(10t) \end{array}\right) \times 10^{-3} \text{N},$$  

$$\vec{r}_{D}^B = \left(\begin{array}{c} 6 \sin(0.5t) + 0.6 \sin(50t + \frac{\pi}{4}) \\ 3 \sin(0.5t + \frac{\pi}{4}) + 0.3 \sin(50t + \frac{\pi}{4}) \\ 4 \sin(0.5t + \frac{\pi}{4}) + 0.4 \sin(50t) \end{array}\right) + \left(\begin{array}{c} -2 \\ 6 \\ 5 \end{array}\right) \times 10^{-3} \text{N} \cdot \text{m}.$$  

<table>
<thead>
<tr>
<th>Inclination (deg)</th>
<th>Argument of Perigee (Deg)</th>
<th>True Anomaly at Time Zero (Deg)</th>
<th>RAAN (Deg)</th>
<th>Eccentricity</th>
<th>Perigee Altitude (km)</th>
</tr>
</thead>
<tbody>
<tr>
<td>45</td>
<td>30</td>
<td>0</td>
<td>50</td>
<td>0</td>
<td>500</td>
</tr>
</tbody>
</table>

Table 2. Orbital parameters of the leader spacecraft.

Figure 2. Block diagram of the simulation.
In Figure 3, the relative position vector histories of the follower spacecraft relative to the desired frame $F_D$ are given. It can be observed that the components of the position vector converge to the neighborhood of zero. The results imply that the follower spacecraft moves to its desired position and stays in the desired position. Figure 4 shows the time histories of the velocity vector of the follower spacecraft relative to the desired coordinate frame. The variations of the relative velocities are consistent with those of the corresponding position histories in Figure 3. Clearly, the convergence of the controller in this paper is smoother and the steady errors smaller than those of the control law in Ref. [13].

![Figure 3](image1.png)

**Figure 3.** Relative position histories of the proposed method and the PD-like method in [13].

![Figure 4](image2.png)

**Figure 4.** Relative velocity histories of the proposed method and the PD-like method in [13].

Figure 5 illustrates the MRP histories of the follower spacecraft relative to the desired coordinate frame. Similar to the results of the translational motion, the attitude errors are smaller as well. Figure 6 provides the profiles of the relative angular velocity, where the relative angular velocity governed by the proposed controller changes more aggressively to enable fast convergence. Obviously, the proposed control scheme outperforms the PD-like controller in attitude control, too.

![Figure 5](image3.png)

**Figure 5.** Relative attitude histories of the proposed method and the PD-like method in [13].
Figure 6. Relative angular velocity histories of the proposed method and the PD-like method in [13].

The time histories of the control force and torque are presented in Figures 7 and 8, respectively. Though the relative position of the finite-time controller converges faster than that in [13], both the force and torque magnitudes of the former are much smaller than those of the latter. Table 3 exhibits more details about the results obtained by comparing the final accuracy of the two methods. The stronger robustness of the proposed approach is indicated by the higher steady accuracy in the presence of disturbances.

Table 3. Steady accuracy comparison of the two control laws.

<table>
<thead>
<tr>
<th>Steady Accuracy of the Method in [13]</th>
<th>Steady Accuracy of the Method in this Paper</th>
</tr>
</thead>
<tbody>
<tr>
<td>$|r_{BD}|_2$ (m)</td>
<td>$1.01 \times 10^{-2}$</td>
</tr>
<tr>
<td>$|v_{BD}|_2$ (m/s)</td>
<td>$9.72 \times 10^{-4}$</td>
</tr>
<tr>
<td>$|p_{BD}|_2$</td>
<td>$4.60 \times 10^{-3}$</td>
</tr>
<tr>
<td>$|\omega_{BD}|_2$ (rad/s)</td>
<td>$5.33 \times 10^{-4}$</td>
</tr>
</tbody>
</table>

Figure 7. Control force profiles of the proposed method and the PD-like method in [13].

Figure 8. Control torque profiles of the proposed method and the PD-like method in [13].

5. Conclusions

For the control problem of 6-DOF leader–following spacecraft formation, the relative position–attitude motion between spacecraft is described by a unified model based on twistors. A finite-time control scheme is derived for integrated pose control with the
application of homogeneous theory. The effectiveness of the control algorithm is manifested by numerical simulations, and the advantages are shown by comparison with an asymptotic control scheme. Designing an integrated pose controller in the framework of twistors is more straightforward than conventional approaches, and the finite-time algorithm achieves stronger robustness than asymptotic ones. In the future, twistor representation can be extended to distributed cooperative control of multi-spacecraft formations and has great potential for obtaining more accurate and robust control protocols.

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**Abbreviations**
The following abbreviations are used in this manuscript:

- APF Artificial potential field
- MRP Modified Rodrigues parameters
- PD Proportion differentiation
- 6-DOF Six-degree-of-freedom
- UAS Uniformly asymptotically stable
- UFTS Uniformly finite-time stable

**References**


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