



Article Enthusiasm and Skepticism: Two Pillars of Science—A Nonextensive Statistics Case

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Abstract: Science and its evolution are based on complex epistemological structures. Two of the pillars of such a construction definitively are enthusiasm and skepticism, both being ingredients without which solid knowledge is hardly achieved and certainly not guaranteed. Our friend and colleague Jean Willy André Cleymans (1944–2021), with his open personality, high and longstanding interest for innovation, and recognized leadership in high-energy physics, constitutes a beautiful example of the former. Recently, Joseph I. Kapusta has generously and laboriously offered an interesting illustration of the latter pillar, in the very same field of physics, concerning the very same theoretical frame, namely, nonextensive statistical mechanics and the nonadditive *q*-entropies on which it is based. I present here a detailed analysis, point by point, of Kapusta's 19 May 2021 talk and, placing the discussion in a sensibly wider and updated perspective, I refute his bold conclusion that indices *q* have no physical foundation.

Keywords: Jean Cleymans; Joseph Kapusta; thermodynamics; nonadditive entropies; nonextensive statistical mechanics; high-energy physics



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1. Reminiscences Related to Jean Cleymans

I had the privilege of personally meeting Jean Cleymans in two occasions. The first one was at CERN (European Organization for Nuclear Research) for the Heavy Ion Forum held in 2013–2014 to discuss *q*-statistics in high-energy physics. His experience and scientific weight among his CERN colleagues were noticeable. That was my first visit to CERN, where I stayed a couple of days. The second occasion was during a visit, over several days, to the University of Cape Town, when I also had the opportunity to lecture and discuss, with him and his international research group, on the foundations and applications of nonadditive entropies and nonextensive statistical mechanics.

In the bibliography in [1], over 9000 articles, by nearly 16,000 authors, are registered on this research area, which started in 1988 [2]. Among them, more than one thousand concern high-energy physics (high-energy collisions, possible connections to quantum chromodynamics (QCD), thermofractals, solar physics, astronomy, astrophysics, black holes, cosmology, dark energy; theoretical, observational, experimental). I could find therein fourty two articles by Cleymans and co-authors [3–44], spanning over one decade. His longstanding *enthusiasm* about the subject, at least in what concerns high-energy physics, can be hardly doubted.

2. Analysis of Joseph Kapusta's Talk "A Primer on Tsallis Statistics for Nuclear and Particle Physics" (19 May 2021)

2.1. Preliminaries

On 19 May 2021, Joseph Kapusta delivered an online talk within the Theoretical Physics Colloquium series that Igor Shovkovy has been hosting at Arizona State University. The talk was entitled "A primer on Tsallis statistics for nuclear and particle physics", and

it is accessible on the Internet [45]. Making it publicly accessible was, in some sense, a fortunate initiative. Indeed, in spite of the fact that my own name appears in the title of the talk, I never received notice of it, rather regretfully. Nevertheless, fortuitously, I could eventually hear, in 27 December 2021, Kapusta's talk when I became aware of its existence through an anonymous referee report on a project involving *q*-concepts.

However, why am I giving all these details? Basically, because I agree with the old dictate "silence gives consent" ("qui tacet consentit"), and by no means can I agree with nor remain silent regarding the final conclusions offered by the speaker of that specific talk. Indeed, in what follows, I intend to offer, essentially point by point, an analysis of and a reply to Kapusta's talk, not only in what concerns the logical implications of his statements and conclusions, but—quite appropriate under the present circumstances—also in what concerns the spirit of the "analyse du discours" [46–48]. This is why I focus here on the talk itself and not on the written version, which Kapusta published some months later [49]. Let me add for completeness that, as soon as I became aware of the existence of that talk on YouTube, I proposed an open academic debate between Kapusta and myself, possibly organized within the same or similar frame of the Arizona State University Colloquia series (or any other appropriate forum in fact), but I did not receive back a signal of operational interest. Unfortunately, an opportunity for a free and direct exchange of scientific ideas focusing on a currently active research area was lost.

Before starting, let me by all means make absolutely clear my genuine gratitude to Kapusta for having dedicated a sensible amount of his time and knowledge to analyze and better understand the theory that I proposed over three decades ago (in [2,50] and elsewhere). It is my understanding that, by so proceeding, Kapusta also attempted to serve science along its best lines, in this case through *skepticism*, as it becomes clear in what follows.

2.2. On the Title of Kapusta's Talk

The choice of the word "primer" in the title is suggestive. For this word one finds in the Merriam-Webster Dictionary:

- A small book for teaching children to read;
- A small introductory book on a subject;
- A short informative piece of writing.

In the Cambridge Dictionary, one finds:

- A small book containing basic facts about a subject, used especially when you are beginning to learn about that subject;
- A basic text for teaching something.

It appears therefore that the word "primer" tends to induce the audience in the sense that basic undeniable facts will be presented, as opposed to opinions or beliefs. In classical philosophical Plato terms, "primer" induces us to think about "episteme" (from the Greek $\epsilon\pi\iota\sigma\tau\eta\mu\eta$, related to facts) and not about "doxa" (from the Greek $\delta\delta\xi\alpha$, opinion). But, as the reader can verify, what one hears in various parts of the talk is a blurring between these two concepts, facts and opinion, a blurring between "text" and "context".

Along a similar vein, the expression "Tsallis and his collaborators" that Kapusta used several times is not exempted from ambiguity. Indeed, there are my collaborators—co-authors, quite frequently—and there are also thousands of scientists (see [1]) who became interested in the subject and have published, along the years, thousands of articles, some of which presenting aspects that I would surely endorse, and others whose statements are certainly not to be confused with my own.

2.3. Nonadditive Entropy S_q : Where It Comes from

In his talk, Kapusta declared a couple of times that he did not know where S_q came from. Although spread in the literature for decades, let us summarize in the present

occasion where it comes from. Like the Boltzmann–Gibbs entropic functional, it is basically a *postulate*, as stated in [2]. There is, however, a simple rationale behind it. Let us describe it.

An entropic functional must be invariant under permutation of states; indeed, what matters are the probabilities $\{p_i\}$ and not the states themselves. The simplest manner which satisfies this property is to involve $\sum_i f(p_i)$, and the simplest form for $f(p_i)$ is just a power-law, i.e., $f(p_i) \propto p_i^q$ ($q \in \mathbb{R}$). Consistently, the entropic functional being constructed is assumed to be of the form $S_q(\{p_i\}) = \Phi(\sum_i p_i^q)$, with $\sum_i p_i = 1$. The simplest form for $\Phi(x)$ is to be linear, i.e., $S_q(\{p_i\}) = A(q) + B(q) \sum_i p_i^q$. Probabilistic certainty should make the entropy to vanish. Therefore, if it is possible for a state j to exist such that $p_i = 1$, then $\sum_i p_i^q = 1$, hence A(q) + B(q) = 0. Consequently, $S_q(\{p_i\}) = A(q)(1 - \sum_i p_i^q)$. Now, the whole intuition for the present postulate, inspired in multifractals as declared in [2], is to provide a *bias* for the state probability. More definitely, to compare p_i^q with p_i for any $p_i \in (0,1)$: $p_i^q > p_i$ if q < 1, $p_i^q < p_i$ if q > 1, and finally $p_i^q = p_i$ if q = 1. As one can see, q = 1 simply means absence of bias, and this is intuitively assumed to well correspond to the Boltzmann–Gibbs (GB) entropy variable, $S_{BG} = k \sum_{i} p_i \ln(1/p_i)$. Therefore, $\lim_{q\to 1} A(q)$ must diverge in order to compensate for the fact that $\lim_{q\to 1} [1 - 1]$ $\sum_{i} p_{i}^{q} = \lim_{q \to 1} \left[1 - \sum_{i} p_{i} p_{i}^{q-1} \right] = \lim_{q \to 1} \left[1 - \sum_{i} p_{i} e^{(q-1) \ln p_{i}} \right] = \lim_{q \to 1} \left\{ 1 - \sum_{i} p_{i} \left[1 + (q - 1) + p_{i} \right] \right\}$ 1) $\ln p_i$] = $\lim_{q \to 1} (q-1) [-\sum_i p_i \ln p_i] = 0$; moreover, A(q) must have the dimensions of the constant k, hence be proportional to k. The simplest such possibility satisfying $\lim_{q\to 1} S_q\{p_i\} = S_{BG}\{p_i\}$ is A(q) = k/(q-1), and we are done. The entropy S_q is therefore defined by

$$S_{q}(\{p_{i}\}) \equiv k \frac{1 - \sum_{i=1}^{W} p_{i}^{q}}{q - 1} = k \sum_{i=1}^{W} p_{i} \ln_{q} \frac{1}{p_{i}}$$
$$= -k \sum_{i=1}^{W} p_{i}^{q} \ln_{q} p_{i} = -k \sum_{i=1}^{W} p_{i} \ln_{2-q} p_{i} \quad (S_{1} = S_{BG}), \quad (1)$$

where the (monotonically increasing) *q*-logarithmic function is defined through

$$\ln_q z \equiv \frac{z^{1-q} - 1}{1-q} \quad (\ln_1 z = \ln z),$$
(2)

with $d \ln_q z/dz = 1/z^q$. The quantity $\sigma(p) \equiv \ln(1/p)$ is referred to as "surprise" [51] or "unexpectedness" [52]. It vanishes when the probability p equals unity and diverges when the probability $p \to 0$. One can consistently define q-surprise or q-unexpectedness as $\sigma_q(p) \equiv \ln_q(1/p)$, hence $\sigma_q(1) = 0$ and $\sigma_q(0)$ attains its maximum (infinity if $q \leq 1$). With this definition, S_q can be rewritten as follows:

$$S_q = k \sum_{i=1}^{W} p_i \ln_q(1/p_i) \equiv k \langle \ln_q(1/p_i) \rangle = k \langle \sigma_q \rangle, \qquad (3)$$

where $\langle ... \rangle$ denotes the mean value. Equation (3) satisfies in fact the most general form for a "trace-form" entropic functional, $S_{\Psi}(\{p_i\}) = k \sum_i p_i \ln_{\Psi}(1/p_i)$, $\ln_{\Psi}(z)$ being a generic monotonically increasing function which satisfies $\ln_{\Psi}(1) = 0$; for details, see [50,53].

In addition, an entropic functional $S(\{p_i\}; \eta)$ is said "composable" if it satisfies, for two *probabilistically independent* systems *A* and *B*, the property $S(A + B)/k = F(S(A)/k, S(B)/k; \eta)$, where η is a set of fixed indices characterizing the functional (e.g., for S_q , it is $\eta \equiv q$); the notation $\{0\}$ is used here to indicate absence of any such index. $F(x, y; \{\eta\})$ is assumed to be a smooth function of (x, y), which depends on a (typically small) set of universal indices $\{\eta\}$, defined in such a way that $F(x, y; \{0\}) = x + y$ (additivity), which corresponds to the Boltzmann–Gibbs entropy. Furthermore, $F(x, y; \{\eta\})$ is assumed to satisfy $F(x, 0; \{\eta\}) = x$ (null-composability), $F(x, y; \{\eta\}) = F(y, x; \{\eta\})$ (symmetry), $F(x, F(y, z; \{\eta\}); \{\eta\}) = F(F(x, y; \{\eta\}), z; \{\eta\})$ (associativity); for details and thermodynamical motivation, see [50,53].

The Enciso–Tempesta theorem [54] proves that S_q is the *unique* entropic functional which simultaneously is a trace form, composable, and includes the Boltzmann–Gibbs form as a particular instance. These properties play a crucial role in the validity, for *all values* of q, of the Einstein requirement for likelihood factorization [55]. This important property is consistent with the fact that the Boltzmann–Gibbs statistical mechanics is *sufficient but not necessary* for the validity of thermodynamics and its Legendre structure.

The word "postulate" used above deserves a comment. For the Boltzmann–Gibbs entropy, it is possible to replace its postulated form by a set of four axioms. That was first shown by Shannon in his celebrated uniqueness theorem [56,57], and later on, in a more elegant form, by Khinchin in his own uniqueness theorem [58,59]. Both theorems have been q-generalized into Santo's uniqueness theorem [60] and Abe's uniqueness theorem [61], respectively. Let us finally mention that, in addition to the above uniqueness theorems (Santos in 1997, Abe in 2000, and Enciso and Tempesta in 2017), S_q is unique in other relevant senses as well, namely, the Topsøe factorizability in game theory [62], the Amari–Ohara–Matsuzoe conformally invariant geometry [63] and the Biro–Barnafoldi–Van thermostat universal independence [64].

2.4. Additivity versus Extensivity

All the way along his talk, Kapusta did not distinguish—clearly enough, and even at all—the concepts of "extensivity" and "additivity", applicable to both entropy and energy. This is quite unfortunate since this distinction ought to be made in any introductory talk on the subject. Indeed, it plays a foundational role in nonextensive statistical mechanics. Let us address now these two important notions, focusing specifically on entropic additivity and entropic extensivity.

Following O. Penrose [65], an entropic functional $S(\{p_i\})$ is said "additive" if, for two *probabilistically independent* systems *A* and *B* (i.e., $p_{ij}^{A+B} = p_i^A p_j^B$), one verifies S(A + B) = S(A) + S(B), in other words, if

$$S(\{p_i^A p_i^B\}) = S(\{p_i^A\}) + S(\{p_i^B\})$$
(4)

is verified.

Otherwise, $S({p_i})$ is said "nonadditive". It immediately follows that S_{BG} is additive. In contrast, S_q satisfies

$$\frac{S_q(A+B)}{k} = \frac{S_q(A)}{k} + \frac{S_q(B)}{k} + (1-q)\frac{S_q(A)}{k}\frac{S_q(B)}{k},$$
(5)

hence

$$S_q(A+B) = S_q(A) + S_q(B) + \frac{1-q}{k} S_q(A) S_q(B).$$
 (6)

Therefore, unless $(1 - q)/k \rightarrow 0$, S_q is nonadditive.

Let us now address the other important entropic concept, namely, extensivity. An entropy S(N) is said "extensive" if a specific entropic functional is applied to a specific class of many-body systems with $N = L^d$ particles, where *L* is its dimensionless linear size and *d* its spatial dimension, and satisfies the thermodynamical expectation

$$0 < \lim_{N \to \infty} \frac{S(N)}{N} < \infty, \tag{7}$$

hence, $S(N) \propto N$ for $N \gg 1$. Therefore, entropic additivity only depends on the entropic functional, whereas entropic extensivity depends on *both* the chosen entropic functional *and* the system itself (i.e., its constituents and the correlations among them).

Let us illustrate this fundamental distinction through four, among infinitely many, equal-probability typical examples of W(N) $(N \to \infty)$, where W is the total number of possibilities whose probability does not vanish.

- Exponential class $W(N) \sim A\mu^N \quad (A > 0; \mu > 1)$: This is the typical case within the Boltzmann–Gibbs theory. One gets: $S_{BG}(N) = k \ln W(N) \sim N \ln \mu + \ln A \propto N$, therefore, S_{BG} is extensive, as thermodynamically required.
- Power-law class $W(N) \sim BN^{\rho}$ $(B > 0; \rho > 0)$: One should not use S_{BG} since it implies $S_{BG}(N) = k \ln W(N) \sim \rho \ln N + \ln B \propto \ln N$, thus violating thermodynamics. One verifies instead that $S_{q=1-1/\rho}(N) = k \ln_{q=1-1/\rho} W(N) \propto N$, as thermodynamically required.
- Stretched exponential class $W(N) \sim C\nu^{N\gamma}$ (C > 0; $\nu > 1$; $0 < \gamma < 1$): In this case, no value of q exists which would yield an extensive entropy S_q . One can instead use $S_{\delta} \equiv k \sum_{i=1}^{W} p_i [\ln 1/p_i]^{\delta}$ [50] with $\delta = 1/\gamma$. Indeed, $S_{\delta=1/\gamma}(N) = k [\ln W(N)]^{1/\gamma} \propto N$, as thermodynamically required.
- Logarithmic class $W(N) \sim D \ln N$ (D > 0): In this case, no values of (q, δ) exist which are able to yield an extensive entropy $S_{q,\delta} \equiv k \sum_i p_i [\ln_q (1/p_i)^{\delta}$ [66]. Instead, one can use the Curado entropy [53], $S_{\lambda}^{C}(N) = k [e^{\lambda W(N)} - e^{\lambda}]$ with $\lambda = 1/D$. Indeed, one can verify that $S_{\lambda=1/D}^{C}(N) \propto N$, as thermodynamically required.

These four paradigmatic classes are described in Figure 1.

SYSTEMS W(N) (equiprobable)	ENTROPY S _{BG} (ADDITIVE)	ENTROPY S_q $(q \neq 1)$ (NONADDITIVE)	ENTROPY S_{δ} ($\delta \neq 1$) (NONADDITIVE)	ENTROPY S_{λ}^{C} ($\lambda > 0$) (NONADDITIVE)
$\sim A \mu^{N} (A > 0, \mu > 1)$	EXTENSIVE	NONEXTENSIVE	NONEXTENSIVE	NONEXTENSIVE
$\sim B N^{\rho} (B > 0, \rho > 0)$	NONEXTENSIVE	EXTENSIVE $(q=1-1/\rho)$	NONEXTENSIVE	NONEXTENSIVE
$\sim Cv^{N^{\gamma}}(C>0, v>1, 0<\gamma<1)$	NONEXTENSIVE	NONEXTENSIVE	EXTENSIVE $(\delta = 1/\gamma)$	NONEXTENSIVE
$\sim D \ln N \ (D > 0)$	NONEXTENSIVE	NONEXTENSIVE	NONEXTENSIVE	EXTENSIVE $(\lambda = 1/D)$

Figure 1. Typical behaviors of W(N) (number of nonzero-probability states of a system with N random variables) in the $N \to \infty$ limit and entropic functionals which, under the assumption of *equal probabilities* for all states with nonzero probability, yield extensive entropies for specific values of the corresponding (nonadditive) entropic indices. Concerning the exponential class $W(N) \sim A\mu^N$, S_{BG} is not the unique entropy that yields entropic extensivity; the (additive) Renyi entropic functional, $S_q^R \equiv k(\ln \sum_i p_i^q)/(1-q)$, also is extensive for all values of q. Analogously, concerning the stretched-exponential class $W(N) \sim Cv^{N^{\gamma}}$, the (nonadditive) entropic functional S_{δ} is not unique. All the entropic families illustrated contain S_{BG} as a particular case, excepting the Curado entropy, S_{λ}^C , which is appropriate for the logarithmic class $W(N) \sim D \ln N$. In the limit $N \to \infty$, the inequalities $\mu^N \gg v^{N^{\gamma}} \gg N^{\rho} \gg \ln N \gg 1$ are satisfied, hence $\lim_{N\to\infty} v^{N^{\gamma}}/\mu^N = \lim_{N\to\infty} N^{\rho}/\mu^N = \lim_{N\to\infty} \ln N/\mu^N = 0$. This exhibits that, in all these nonadditive cases, the occupancy of the full phase space corresponds essentially to a zero Lebesgue measure, similarly to a whole class of (multi)fractals. If the equal-probabilities hypothesis is not satisfied, a specific analysis becomes necessary and the results might be different. Taken from [50].

At this point, it is pertinent to remind Einstein's 1949 comment [67]: "A theory is the more impressive the greater the simplicity of its premises is, the more different kinds of

things it relates, and the more extended is its area of applicability. Therefore, the deep impression that classical thermodynamics made upon me. It is the only physical theory of universal content concerning which I am convinced that, within the framework of applicability of its basic concepts, it will never be overthrown".

To better understand the strength of these words, a metaphor can be used. Within Newtonian mechanics, there is the known Galilean composition of velocities, $v_{13} = v_{12} + v_{13} + v_{13}$ v_{23} . In special relativity, this law is generalized into $v_{13} = \frac{v_{12} + v_{23}}{1 + v_{12}v_{23}/c^2}$, where *c* is the speed of light, thus violating additivity. Why did Einstein abandon the simple linear composition of Galileo? Because he had a higher goal, namely, to unify mechanics and Maxwell's electromagnetism, and, for this, he had to impose the invariance with regard to the Lorentz transformation. One can, thus, see the violation of the linear Galilean composition as a small price to pay for a major endeavor. Similarly, what is expressed in Figure 1, is that q-generalizing the *linear* composition law of S_{BG} , with regard to independent systems, into the nonlinear composition (5) may be seen as a small price to pay for a major endeavor, namely, to always satisfy the Legendre structure of thermodynamics. However, it is mandatory to register here that such viewpoint is nevertheless not free from controversy, in spite of its simplicity. For example, the known expression of Bekenstein and Hawking for the entropy for a black hole is proportional to its surface instead of to its volume, therefore violating the above requirement *if* the black-hole is assumed to be a three-dimensional object; see [66].

Before leaving this point, it is fair to mention that Kapusta's lack of distinction between entropic extensivity and entropic additivity would not be particularly surprising, were it not for his own perception of his talk as a "primer". In any case, such confusion constitutes, still today, no exception in the literature even if my co-author Murray Gell-Mann and myself started around 2003 to emphasize the distinction in my book [50] and in our paper [68].

2.5. On the Nature of the Constraints Used for Entropic Optimization

The relevance of the invariance, under a zero-energy shift at *fixed* inverse effective temperature β_q , of the distribution of probabilities extremizing S_q took, from [2] through [69] to [70], close to ten years to become clarified. As pedagogically explained by Kapusta, this process led, through types I and II, to the type III or third path, which was specially focused in his talk. The equivalence, under appropriate transformations of q and, consistently, of the Lagrange parameters (α, β) , associated with these three paths was later on established in [71,72]. Indeed, the possibility of linking S_{2-q} (instead of S_q) with the extremizing q-exponential distribution and *linear* constraints such as $\sum_i p_i E_i = \langle E \rangle$ (instead of the *nonlinear* constraint $\sum_i p_i^q E_i / \sum_i p_i^q = \langle E \rangle_q$) was later on used in [73–82] on the basis of identities (1), and has nowadays become quite popular due to its operational simplicity. This link is based on a nonlinear Fokker–Planck Equation (with nonlinear exponent (2 – q)) in the presence of a confining potential [83,84]. This equation has, as its basic stationary state, a q-Gaussian and not a (2 – q)-Gaussian. Further examples exhibiting the same type of duality are available at [85,86].

Let us mention at this stage that several central related issues were regretfully ignored in Kapusta's talk, namely:

• The peculiar way the nonlinear constraint of type III was phrased by the speaker left floating in the audience (which even repeated this later on, with the tacit agreement of the speaker) that this assumption was violating the theory of probabilities. There is *no* such thing.

Indeed, the above *nonlinear* constraint in p_i is completely equivalent to a *linear* constraint expressed in terms of $\sum_i P_i^{(q)} E_i$ where $P_i^{(q)} \equiv p_i^q / \sum_j p_j^q$ (hence $p_i = (P_i^{(q)})^{1/q} / \sum_j (P_j^{(q)})^{1/q}$, with $\sum_i p_i = \sum_i P_i^{(q)} = 1$) is the so-called "escort distribution," defined in the theory of probabilities; see [87] and references therein.

• The requirement for system-independence for the adopted extremization procedure consists of having for both the norm constraint, expressed in terms of $\sum_i p_i$, and the

energy constraint, expressed in terms of $\sum_i P_i^{(q)} E_i$, one and the same upper admissible value for q. Let us illustrate this through an example. If the set of eigenvalues $\{E_i\}$ are nondegenerate, the extremizing q-exponential distribution with q > 1 asymptotically behaves as $1/E_i^{1/(q-1)}$. Therefore, its norm is well defined up to q = 2. The same happens with the constraint $\sum_i P_i^{(q)} E_i$. Indeed, it asymptotically behaves as $E_i/E_i^{q/(q-1)} = 1/E_i^{1/(q-1)}$, hence it has the same upper bound admissibility, i.e., once again q = 2. In strong variance, if one were to use here the usual $\sum_i p_i E_i$ constraint, the asymptotic behavior would be given by $E_i/E_i^{1/(q-1)}$, hence its upper admissible value would be q = 3/2, which *differs* from the norm admissibility value, q = 2. The general importance of this point is lengthily discussed in [88].

- In the *q*-generalized central limit theorem and Lévy–Gnedenko's limit theorem [89,90], the escort mean values emerge naturally from the mathematical operations themselves.
- All mean values must be mathematically defined for the fat-tailed *q* > 1 distributions. This does happen with the appropriate escort averages but fails with the normal linear averages. A simple numerical example with *q*-Gaussians and its astonishing practical consequences are illustrated in [91].

2.6. On Ad Hoc Constraints for Optimizing the Entropy

Kapusta referred, in a kind of appreciative style, to a 2004 argument by Zanette and Montemurro [92], which would "disqualify" S_q as a "physical" entropy. It refers to the trivial mathematical fact that by adjusting in an ad hoc manner the constraints under which the entropy optimization is to be done, one can obtain virtually any desired distribution. This mathematical feature applies to all admissible entropic functionals (including S_{BG}) and has been an overcome issue for a long time. Indeed, already in 1983, Montroll and Shlesinger [93] exhibited how the nontrivial constraint can be adapted in order to obtain Lévy distributions from the *classical Boltzmann–Gibbs* entropy. The author's epistemological dissatisfaction is evident in the paper. Indeed, one is forced to *first* know the distribution that is aimed, and *then*, through a sort of reverse engineering, write down the nontrivial constraint in order to get it. This is actually possible for any entropic functional, as mentioned above, but provides no useful information at all. Constraints within information theory are not to be freely manipulated: they must reflect constants of motion and similar quantities. For a simple real random variable *x*, they must restrict to robust constraints such as the values of $\langle x \rangle$ and/or of $\langle x^2 \rangle$. It is up to the entropic functional, and *not* to the ad hoc constraints, to do the most meaningful and nontrivial job.

We argue here that, since it is known to be so for *any* entropic form and, in particular, for the (additive) Boltzmann–Gibbs entropy, S_{BG} (see [94]), the critique brings absolutely no novelty to the area. In other words, it has *nothing* special to do with the entropy S_q . In favor of the constraints based on the usual simple variables (typically, the averaging of the random variable x_i or of x_i^2 , where x_i is to be identified according to the nature of the system), we argue, and this for *all* entropic forms, that they can hardly be considered as arbitrary, as the authors of [92] seemingly consider. Indeed, once the natural variables of the system have been identified (e.g., constants of motion of the system, such as the energy for Hamiltonian systems), the variable itself and, in some occasions, its square are the most basic quantities to be constrained. Such constraints are used in hundreds (perhaps thousands) of useful applications outside (and also inside) thermodynamical systems, along the information theory lines of Jaynes and Shannon, and more recently of A. Plastino and others. Furthermore, this is so for S_{BG} , S_q , and any other entropic form. Rebuttals of the Zanette and Montemuro outdated criticism can be found in [50,95,96].

As a final comment, let us mention that statistical mechanics is much more that just a stationary-state (e.g., thermal equilibrium) distribution. Indeed, under exactly the same constraints, the optimization of S_{BG} and $(S_{BG})^3$ yields the same distribution. This is not a sufficient reason for using $(S_{BG})^3$ instead of S_{BG} in a thermostatistical theory which must

also satisfy various thermodynamical requirements, including its Legendre-transformation structure.

2.7. Thermodynamics and Legendre Transformations

Before focusing on the Legendre structure itself, let us review some long-known facts concerning long-range interactions. Let us consider a *d*-dimensional classical many-body system with, e.g., attractive two-body isotropic interactions decaying with a dimensionless distance $r \ge 1$ as $-A/r^{\alpha}$ (A > 0, $\alpha \ge 0$), and with an infinitely repulsive potential for $0 \le r \le 1$. At zero temperature *T*, the total kinetic energy vanishes, and the potential energy per particle is proportional to $\int_{1}^{\infty} dr r^{d}r^{-\alpha}$. This quantity converges if $\alpha/d > 1$ and diverges otherwise. These two regimes are from now on referred to as "short-range" and "long-range" interactions, respectively; see Figure 2.

Furthermore, they can be shown to, respectively, correspond, within the Boltzmann-Gibbs theory, to *finite* and *divergent* partition functions. This is indeed the point that was addressed by Gibbs himself [97]: "In treating of the canonical distribution, we shall always suppose the multiple integral in equation (92) [the partition function, as it is called nowadays] to have a finite value, as otherwise the coefficient of probability vanishes, and the law of distribution becomes illusory. This will exclude certain cases, but not such apparently, as will affect the value of our results with respect to their bearing on thermodynamics. It will exclude, for instance, cases in which the system or parts of it can be distributed in unlimited space [...]. It also excludes many cases in which the energy can decrease without limit, as when the system contains material points which attract one another inversely as the squares of their distances. [...]. For the purposes of a general discussion, it is sufficient to call attention to the assumption implicitly involved in the formula (92)".



Figure 2. The so-called "extensive systems" ($\alpha/d > 1$ for the classical ones with the exponent α and dimension *d*) correspond to an extensive total energy and typically involve *absolutely convergent* series, whereas the so-called "nonextensive systems" ($0 \le \alpha/d < 1$ for the classical ones) correspond to a superextensive total energy and typically involve *divergent* series. The marginal systems ($\alpha/d = 1$ here) typically involve *conditionally convergent* series, which therefore depend on the boundary conditions, i.e., typically on the external shape of the system. Capacitors constitute a notorious example of the $\alpha/d = 1$ case. The standard Lennard–Jones gas is located at (d, α) = (3,6), thus belonging to the extensive class of systems. Taken from [53].

In a vein slightly differing from the standard Boltzmann–Gibbs recipe, which would demand integration up to infinity in $\int_1^\infty dr r^d r^{-\alpha}$, let us assume that the *N*-particle system

is roughly homogeneously distributed within a limited sphere. Then, the potential energy per particle scales as follows:

 $\frac{U_{\text{pot}}(N)}{N} \propto -A \int_{1}^{N^{1/d}} dr \, r^{d-1} \, r^{-\alpha} = -\frac{A}{d} N^* \,, \tag{8}$

with

$$N^{\star} \equiv \frac{N^{1-\alpha/d} - 1}{1 - \alpha/d} = \ln_{\alpha/d} N \sim \begin{cases} \frac{1}{\alpha/d - 1} & \text{if } \alpha/d > 1; \\ \ln N & \text{if } \alpha/d = 1; \\ \frac{N^{1-\alpha/d}}{1 - \alpha/d} & \text{if } 0 \le \alpha/d < 1. \end{cases}$$
(9)

Therefore, in the $N \to \infty$ limit, $U_{pot}(N)/N$ approaches a constant ($\propto -A/(\alpha - d)$) if $\alpha/d > 1$, and diverges like $N^{1-\alpha/d}/(1-\alpha/d)$ if $0 \le \alpha/d < 1$ (it diverges logarithmically if $\alpha/d = 1$). In other words, the total potential energy is extensive for short-range interactions ($\alpha/d > 1$), and nonextensive for long-range interactions ($0 \le \alpha/d \le 1$). Satisfactorily enough, Equation (9) recovers the characterization with $\int_1^{\infty} dr r^d r^{-\alpha}$ in the limit $N \to \infty$, but it has the great advantage of providing, for *finite* N, a *finite* value. This fact is now shown to enable us to properly scale the macroscopic quantities in the thermodynamic limit ($N \to \infty$), for *all* values of $\alpha/d \ge 0$ and not only for $\alpha/d > 1$ (as in textbooks of thermodynamics, even if the point is too frequently not made explicitly).

The mathematical structure of classical thermodynamics is based on the Legendre transforms. It is not sufficiently realized that thermodynamics does not depend on microscopic details *only for short-range* interactions. As is illustrated here below, it does depend on quantities such as (α, d) for long-range interactions. Quoting Landsberg [98]: "The presence of long-range forces causes important amendments to thermodynamics, some of which are not fully investigated as yet".

Let us consider a *d*-dimensional homogeneous and isotropic classical fluid constituted by magnetic particles in thermodynamical equilibrium. Its Gibbs free energy is then:

$$G(N, T, p, \mu, H) = U(N, T, p, \mu, H) - TS(N, T, p, \mu, H) + pV(N, T, p, \mu, H) - \mu N - HM(N, T, p, \mu, H),$$
(10)

where (T, p, μ, H) correspond, respectively, to the temperature, pressure, chemical potential and external magnetic field, *U* is the internal energy, *S* is the entropy, *V* is the volume, *N* is the number of particles, and *M* the magnetization.

If the interactions are short-ranged, i.e., if $\alpha/d > 1$, one can divide this equation by N and then take the $N \rightarrow \infty$ limit. One straightforwardly obtains:

$$g(T, p, \mu, H) = u(T, p, \mu, H) - Ts(T, p, \mu, H) + pv(T, p, \mu, H) - \mu - Hm(T, p, H), \quad (11)$$

where $g(T, p, \mu, H) \equiv \lim_{N \to \infty} G(N, T, p, \mu, H) / N$, and analogously for the other variables of the equation.

If the interactions are long-ranged instead, i.e., if $0 \le \alpha/d \le 1$, all the terms of expression (11) diverge, hence, thermodynamically speaking, they are nonsense. Consequently, the generically correct procedure for all values of $\alpha/d \ge 0$, must conform to the following lines:

$$\lim_{N \to \infty} \frac{G(N, T, p, \mu, H)}{NN^{\star}} = \lim_{N \to \infty} \frac{U(N, T, p, \mu, H)}{NN^{\star}} - \lim_{N \to \infty} \left[\frac{T}{N^{\star}} \frac{S(N, T, p, \mu, H)}{N} \right] + \lim_{N \to \infty} \left[\frac{p}{N^{\star}} \frac{V(N, T, p, \mu, H)}{N} \right] - \lim_{N \to \infty} \frac{\mu}{N^{\star}} - \lim_{N \to \infty} \left[\frac{H}{N^{\star}} \frac{M(N, T, p, \mu, H)}{N} \right], \quad (12)$$

hence,

$$g(T^{\star}, p^{\star}, \mu^{\star}, H^{\star}) = u(T^{\star}, p^{\star}, \mu^{\star}, H^{\star}) - T^{\star}s(T^{\star}, p^{\star}, \mu^{\star}, H^{\star}) + p^{\star}v(T^{\star}, p^{\star}, \mu^{\star}, H^{\star}) - \mu^{\star} - H^{\star}m(T^{\star}, p^{\star}, \mu^{\star}, H^{\star}), \quad (13)$$

where the definitions of T^* and of all the other variables are self-explanatory (e.g., $T^* \equiv \lim_{N\to\infty} [T/N^*]$ and $s(T^*, p^*, \mu^*, H^*) \equiv \lim_{N\to\infty} [S(N, T, p, \mu, H)/N]$). In other words, in order to have *finite* thermodynamic equations of states, one must, in general, express them in the (T^*, p^*, μ^*, H^*) variables. If $\alpha/d > 1$, this procedure recovers the usual equations of states, and the usual *extensive* (G, U, S, V, N, M) and *intensive* (T, p, μ, H) thermodynamic variables. However, if $0 \le \alpha/d \le 1$, the situation is more complex, and *three* instead of the traditional *two* classes of thermodynamic variables emerge. One may call them *extensive* (S, V, N, M), *pseudoextensive* (G, U) (superextensive in the present case) and *pseudointensive* (T, p, μ, H) (superintensive in the present case) variables. All the energy-type thermodynamical variables (G, F, U), F being the Helmholtz free energy, give rise to pseudoextensive ones, whereas those which appear in the usual Legendre thermodynamical pairs give rise to pseudointensive ones (T, p, μ, H) and extensive ones (S, V, N, M). Let us emphasize that (S, V, N, M) are extensive in *all* cases; see Figure 3. The exactness of all these scalings has been repeatedly verified in the literature for diverse concrete physical systems, such as fluids, magnets, polymers and percolation; see [50] and references therein.



Figure 3. Representation of the different scaling regimes for classical *d*-dimensional systems. For attractive long-range interactions (i.e., $0 \le \alpha/d \le 1$, α characterizing the interaction range in a potential with the form $1/r^{\alpha}$; for example, Newtonian gravitation corresponds to $(d, \alpha) = (3, 1)$) one can distinguish *three* classes of thermodynamic variables, namely: those, scaling with L^{θ} , named "pseudointensive" (*L* is a characteristic linear length, θ depends on the class of systems); those, scaling with $L^{d+\theta}$ with $\theta = d - \alpha$, the pseudoextensive ones (the energies); and those, scaling with L^{d} (which are always extensive). For short-range interactions (i.e., $\alpha > d$) one has $\theta = 0$ and the energies recover their standard L^{d} extensive scaling, falling within the same class of *S*, *N*, *V*, *M*, etc. (see text for notations of the thermodynamic variables), whereas the previous pseudointensive variables become truly intensive ones (independent of *L*); this is the region, with only *two* classes of variables, that is covered by the traditional textbooks on thermodynamics. For more details, see [50,53,66,99–101].

Let us also emphasize that, consistently:

- The ratio of any two pseudointensive variables (*T*, *p*, μ, *H*,...), e.g., *p*/*T*, is intensive in all cases;
- The ratio of any pseudoextensive variable (*G*, *F*, *U*) with any pseudointensive variable, e.g., *U*/*T*, is extensive in all cases;
- A most important implication is that, in expressions such as $e_q^{-\beta_q \mathcal{H}_N}$ where \mathcal{H}_N is an *N*-body Hamiltonian, the argument $\beta_q \mathcal{H}_N$ is extensive in all cases. This plays a crucial role in the possible *q*-generalization of what is currently referred to as the large

deviation theory. Indeed, the extensivity of $\beta_q \mathcal{H}_N$ appears to mirror, in all cases, the extensivity of the total entropy involved in $r_q N$, r_q being the ratio function (defined within the Large Deviation probability $P(N) = P(0)e_q^{-r_q N}$), seemingly always related to some relative nonadditive entropy *per particle* [102–106].

2.8. Boltzmann Equation

Kapusta also focused on the current Boltzmann equation associated with the $a + b \leftrightarrow c + d$ reaction. He specifically presented and discussed

$$\frac{df_a}{dt} = \int \frac{d^3 p_b}{(2\pi)^3} \frac{d^3 p_c}{(2\pi)^3} \frac{d^3 p_d}{(2\pi)^3} \left\{ \frac{1}{1+\delta_{cd}} W(c+d\to a+b) f_c f_d \times [1+(-1)^{2s_a} f_a] [1+(-1)^{2s_b} f_b] - \frac{1}{1+\delta_{ab}} W(a+b\to c+d) f_a f_b \times [1+(-1)^{2s_c} f_c] [1+(-1)^{2s_d} f_d] \right\},$$
(14)

where s_i and p_i are the spin and 3-momentum of the *i*th particle, *t* is the time, and δ_{ij} is the Kronecker delta, where i = a, b, c, and *d*. In thermal equilibrium, $df_a/dt = 0$, f_a has the usual Boltzmann–Gibbs expression, and the energy conservation, $\omega_a + \omega_b = \omega_c + \omega_d$ is guaranteed, both in the quantum and the classical cases. In the classical limit, one finds the usual expression, $f = e^{-\beta\omega}$.

Kapusta conveniently pointed out that replacing, in the Boltzmann equation, the traditional molecular chaos hypothesis (*stosszahlansatz*) factorization, $f_c f_d$ by $[f_c^{1-q} + f_d^{1-q} - 1]^{1/(1-q)}$ (and analogously for $f_a f_b$) [107,108], leads to the *q*-exponential distribution as the stationary state, also satisfying simultaneously the energy conservation, $\omega_a + \omega_b = \omega_c + \omega_d$. However, he shared that, essentially, he did not know where that specific replacement came from.

Ulrich Heinz, in the audience, shared with friendly complicity that "After so many papers that I have rejected on Tsallis statistics, I am really happy to hear your talk". About that, not having been communicated in any way that this talk was going to happen online so that I could participate myself, I have no other comment than to say that it is allowed to think that, maybe, not all those rejections were justified. On the other hand, interestingly enough, Heinz also shared that he could, however, "live" with a properly generalized Boltzmann equation.

Let us clarify, at this point, that $[f_c^{1-q} + f_d^{1-q} - 1]^{1/(1-q)}$ is known to correspond, for $q \neq 1$, to a specific (and apparently not rare at all in nature) class of strong correlations between the relevant random variables.

The product *xy* of two real numbers has been conveniently generalized as the following *q*-product [109,110]:

$$x \otimes_{q} y \equiv e_{q}^{\ln_{q} x + \ln_{q} y} = \left[x^{1-q} + y^{1-q} - 1 \right]_{+}^{\frac{1}{1-q}} \quad (x \ge 0, \ y \ge 0) ,$$
(15)

where $[...]_+ = [...]$ if [...] > 0 and vanishes otherwise. Let us list some of its main properties:

• It recovers the standard product as the q = 1 particular instance, i.e.,

$$x \otimes_1 y = xy; \tag{16}$$

• It is commutative for all values of *q*, i.e.,

$$x \otimes_q y = y \otimes_q x; \tag{17}$$

• It is additive under *q*-logarithm for all values of *q*, i.e.,

$$\ln_q(x \otimes_q y) = \ln_q x + \ln_q y, \qquad (18)$$

(referred to as "extensivity"), whereas let us remind that

$$\ln_q(xy) = \ln_q x + \ln_q y + (1-q)(\ln_q x)(\ln_q y)$$
(19)

(referred to as "nonadditivity"). Consistently,

$$e_q^x \otimes_q e_q^y = e_q^{x+y} , \qquad (20)$$

whereas

$$e_q^x e_q^y = e_q^{x+y+(1-q)xy};$$
 (21)

• It is associative for all values of *q*, i.e.,

$$x \otimes_q (y \otimes_q z) = (x \otimes_q y) \otimes_q z = x \otimes_q y \otimes_q z = (x^{1-q} + y^{1-q} + z^{1-q} - 2)^{1/(1-q)};$$
(22)

• It admits unity for all values of *q*, i.e.,

$$x \otimes_q 1 = x;$$
 (23)

• It admits zero under certain conditions, namely

$$x \otimes_{q} 0 = \begin{cases} 0 & \text{if } (q \ge 1 \text{ and } x \ge 0) \text{ or if } (q < 1 \text{ and } 0 \le x \le 1), \\ \\ (x^{1-q} - 1)^{\frac{1}{1-q}} & \text{if } q < 1 \text{ and } x > 1; \end{cases}$$
(24)

 It is distributive with regard to the generalized associative sum analyzed in [111], namely,

$$x \oplus^{(q)} y \equiv e_q^{\ln\left[e^{\ln_q x} + e^{\ln_q y}\right]} = \left\{ 1 + (1-q)\ln\left[e^{\frac{x^{1-q}-1}{1-q}} + e^{\frac{y^{1-q}-1}{1-q}}\right] \right\}^{1/(1-q)}.$$
 (25)

In other words, the *q*-product (15) satisfies all the standard requirements of a full algebraic structure.

Consistently, on the basis of this product, it is possible to generalize, for $q \ge 1$, the Fourier transform of a non-negative function f(x) as follows [89,90]:

$$F_q[f](\xi) \equiv \int dx \, e_q^{i\,\xi\,x} \otimes_q f(x) F_q[f](\xi) = \int_{-\infty}^{\infty} dx \, e_q^{i\,\xi\,x[f(x)]^{q-1}} f(x) \,. \tag{26}$$

It is transparent that this transformation is, for $q \neq 1$, *nonlinear*. Indeed, if $f(x) \rightarrow \lambda f(x)$, λ being any constant, one verifies that $F_q[\lambda f](\xi) \neq \lambda F_q[f](\xi)$. This generalization of the standard Fourier transform $F_1[f](\xi)$ was introduced in order to have a remarkable *property*: it transforms *q*-Gaussians into *q*-Gaussians. Indeed, one verifies:

$$F_{q}\left[B_{q}\sqrt{\beta}e_{q}^{-\beta x^{2}}\right](\xi) = e_{q_{1}}^{-\beta_{1}\xi^{2}},$$
(27)

where

$$(q_1, \beta_1) = \left(\frac{1+q}{3-q}, \frac{3-q}{8\beta^{2-q}B_q^{2(q-1)}}\right) \quad (1 \le q < 2),$$
(28)

 B_q being an appropriate normalizing quantity. Within this frame, and others as well, the central limit theorem has been generalized [89,90], showing that, while averaging a large number of random variables within a wide class of *nonlocal correlations* (yet only partially explored), *q*-Gaussians emerge as attractors in the space of distributions. This provides an epistemological basis for understanding why there are so many *q*-Gaussians (and, consistently, so many *q*-exponentials) in nature; for more details and proofs, see [112]. Moreover, in what concerns the physical significance of the *q*-product (15) in relation with

the Boltzmann transport equation, detailed arguments along a different line are presented as well in [113].

Another relevant line of mathematical research concerns the relations within various sets of *q*-indices (e.g., *q*-triplets) that very frequently emerge in complex systems [68,114].

Let us also mention that, in the context of relaxation times related to the Boltzmann equation, Heinz inquired about an interesting question, namely, whether there is any argument yielding *q*-statistics as an outcome of corresponding dynamics. On general grounds, this not an easy question. However, it is surely useful to have in mind that, in the same way the solution of $dx/dt = -x/\tau$ (with x(0) = 1) is $x = e^{-t/\tau}$, the solution of $dx/dt = -x^q/\tau_q$ (with x(0) = 1) is $x = e_q^{-t/\tau_q}$ [115].

At this point, it might be useful to mention that possible *q*-generalizations of the connections with the Kadanoff–Baym approach have been advanced in [116,117]. Finally, a first step focusing on the Bogoliubov–Born–Green–Kirkwood–Yvon (BBGKY) hierarchy has been introduced in [118].

2.9. On the Second Principle of Thermodynamics

Masoud Shokri, in the audience, inquired about the validity of the second principle for $q \neq 1$. By all means, it appears to be valid in the same way as for the Boltzmann–Gibbs case (q = 1). The following (nonexhaustive) list of arguments that are available in the literature can be mentioned:

- It has been proved that a detailed balance implies the time irreversibility of S_a [119].
- The celebrated Clausius inequality, $\delta Q/T \le dS$, is consistent with the second principle. It has been shown [120] that it remains valid as it stands for q > 1 as well.
- The validity of the *H*-theorem has been proved for a wide class of classical *d*-dimensional many-body overdamped systems including repulsive short-range two-body interactions [73–82].
- The *H*-theorem has been proved for S_q by imposing Galilean-invariance on a Boltzmann lattice model which discretizes the Navier–Stokes equation. The unique value of *q* is determined by the structure of the Bravais lattice that is being focused on; for more details, see [121,122].
- The time evolution of the entropy for low-dimensional maps, or, actually, the Pesin identities, do certainly *not* prove the validity of the second principle. However, the fact that the behavior is quite similar (see [123–125] and references therein) for, e.g., the logistic map at the most chaotic value of its external parameter (with q = 1) and at the Feigenbaum point (with q = 0.2445...) does provide a suggestive indication.

2.10. On the Zeroth Principle of Thermodynamics

Boris Tomasik, in the audience, inquired about the concept of thermal contact among systems in equilibrium, thus tacitly about its transitivity. This highly interesting issue points essentially towards the content of the zeroth principle.

In various overdamped classical *d*-dimensional many-body systems including shortrange repulsive interactions, it has been possible to analytically calculate and numerically verify the validity of *q*-statistics, focusing on space and velocity distributions, equations of states, Carnot cycle, *H*-theorem, effective temperature, *and* the zeroth principle of thermodynamics [73–82]. For example, for repulsive interactions proportional to $1/r^{\alpha}$, it can be proved [82] that the distribution of positions is a *q*-Gaussian with $q = 1 - \alpha/d < 1$, which, in the limit $\alpha/d \rightarrow 1$, recovers the type-II superconductor result q = 0 [73–75]. This class of systems constitute a rare case where nearly all *q*-thermostatistical quantities can be analytically calculated.

However, this may be considered as the infancy of a general proof of the validity of the zeroth principle. Its general discussion remains to be done, and it is related, within *q*-statistics, to the concept of "effective temperature" itself. Indeed, in the above overdamped systems, two different temperatures coexist, namely, the *kinetic* and *effective* temperatures. The former is based on the distributions of velocities, whereas the latter is based on the

spatial distribution. This is in fact quite analogous to what happens in granular packing, sandpiles and the like, where one is forced to distinguish the kinetic temperature from the structural one [126]. This is in variance with Kapusta's repeated expression referring to "the true temperature", without really explaining in physically neat and indisputable terms what "temperature" (or "inverse β ") he is referring to. It is allowed to think that he only has in his mind the usual kinetic concept of temperature, thus ignoring that, in many complex systems, more than one temperature typically emerges; see, e.g., [73–82].

2.11. Indices q: First-Principle Characterization of Universality Classes, or Merely Efficient Fitting *Parameters*?

Kapusta's ultimate question consists of establishing whether or not the indices *q* have a first-principle physical basis. He eventually answers by the negative. Let us point out here that a definitively opposite perspective emerges through the (nonexhaustive) list of counter-examples, either from first principles or at a clear-cut mesoscopic level, in diverse complex systems, including high-energy collisions. Let us emphasize that it is well understood that the analytical determination of the *q* values is, in most of the cases—but fortunately not all!—mathematically intractable.

• First-principle quantum field theory calculation of q for a one-dimensional quantum many-body Hamiltonian system: One such example is the quantum phase transition at T = 0 criticality for the d = 1 first-neighbor Ising ferromagnet in the presence of a transverse magnetic field, and similar quantum systems characterized by a central charge, $c \ge 0$. A subsystem of large linear size L of an infinitely long such chain satisfies $S_{BG}(L) \propto \ln L$, which is nonextensive and therefore violates the Legendre structure of thermodynamics. It turns out, however, that a unique value of q exists such that $S_q(L) \propto L$, thus satisfying thermodynamics. Its value is given by [127]

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$$\eta = \frac{\sqrt{9+c^2}-3}{c},$$
 (29)

which is depicted in Figure 4. As one can see, *q* characterizes here the thermodynamically admissible entropies associated with the criticality universality classes of many-body systems, defined by the central charge *c*.



Figure 4. The index *q* as a function of the inverse central charge, 1/c. The universality classes of some specific models are indicated; for example, c = 1/2 corresponds to the universality class of the short-range finite-spin Ising ferromagnet in the presence of a (critical) transverse field, at temperature T = 0. The Boltzmann–Gibbs value q = 1 is recovered in the $c \rightarrow \infty$ limit. Taken from [55].

Zero-Lebesgue occupancy of the probability space in the *N* → ∞ limit of a system with strongly correlated random binary variables [68]: it is verified that *S_q* is extensive for a unique value of *q*, namely,

$$q = 1 - 1/d$$
, (30)

where *d* is the width, $\forall N$, of the strip with nonvanishing probabilities.

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• Overdamped *d*-dimensional many-body system with short-range repulsive interactions decaying as $1/r^{\alpha}$ ($\alpha \ge d$): The space attractor is a *q*-Gaussian with [82]

$$q = 1 - \alpha/d \,. \tag{31}$$

The limit $\alpha/d = 1$ recovers the superconductor type-II space distribution, which has been proved to correspond to q = 0 [73].

Nonlinear dynamical systems exhibit various direct connections with the time evolution of the entropy and with its consequences. There are, in this respect, two important classes of chaotic behavior, namely: *strong chaos*, characterized by *exponential* sensitivity to the initial conditions (referred to, for classical systems, as having a *positive* maximal Lyapunov exponent); and *weak chaos*, characterized by *subexponential* (frequently, power-law) sensitivity to the initial conditions (referred to, for classical systems, as having a *vanishing* maximal Lyapunov exponent). Let us focus here on an important issue, namely, the time evolution of the entropy while exploring the system's phase space. Let us illustrate this issue with a paradigmatic *dissipative* system, namely, the logistic map, $x_{t+1} = 1 - ax_t^2$ ($x_t \in [-1,1]$; $a \in [0,2]$; t = 0,1,2,...). For a = 2, the system is strongly chaotic since the sensitivity to the initial conditions satisfies $\xi(t) \equiv \lim_{\Delta x(0) \to 0} \Delta x(t) / \Delta x(0) = e^{\lambda t}$ with a Lyapunov exponent $\lambda = \ln 2 > 0$. Consistently, if one starts at t = 0 from a set of M initial conditions in an arbitrarily chosen single window with W of them equally partitioning the interval [-1, 1], one obtains the entropy production per unit time,

$$K_{\rm BG} \equiv \lim_{t \to \infty} \lim_{W \to \infty} \lim_{M \to \infty} \frac{S_{\rm BG}(t)/k}{t} = \lambda > 0, \qquad (32)$$

thus verifying the Pesin identity. One consistently verifies that, for q > 1, $K_q \equiv \lim_{t\to\infty} \lim_{W\to\infty} \lim_{M\to\infty} \frac{S_q(t)/k}{t}$ vanishes, whereas, for q < 1, it diverges. In other words, q = 1 is the unique value of the index for which $S_q(t)$ asymptotically increases *linearly* with time.

If the same operations are applied at the edge of chaos corresponding to the doublebifurcation accumulation point (weak chaos) of the *z*-logistic map ($x_{t+1} = 1 - a|x_t|^z$ with z > 0) and its topologically equivalent one-dimensional dissipative maps, one verifies that $\xi = e_q^{\lambda_q t}$ with $\lambda_q = 1/(1-q)$ [125]. It follows that, for $t \to \infty$, ξ diverges subexponentially, namely as the power-law $\xi \propto t^{1/(1-q)}$. Moreover, it can be shown, through scaling arguments [128], that

$$\frac{1}{1-q} = \frac{1}{\alpha_{\min}} - \frac{1}{\alpha_{\max}},$$
 (33)

where α_{\min} and α_{\max} refer to the (concave) multifractal function $f(\alpha)$, defined in the interval $[\alpha_{\min}, \alpha_{\max}]$ with $f(\alpha_{\min}) = f(\alpha_{\max}) = 0$ [87].

If, for the standard logistic map (z = 2), one actually focuses on the Feigenbaum point (also referred to as the Feigenbaum–Coullet–Tresser point) $a = a_c \equiv 1.40115518909...$

which corresponds to a vanishing Lyapunov exponent λ (i.e., weak chaos), one does verify that $\xi = e_q^{\lambda_q t}$ with $\lambda_q = 1/(1-q)$ [125], as well as

$$\frac{1}{1-q} = \frac{1}{\alpha_{\min}} - \frac{1}{\alpha_{\max}} = \frac{\ln \alpha_F}{\ln 2} ,$$
 (34)

where α_F is the so called Feigenbaum universal constant. From this relation, it follows that

$$q = 0.244487701341282066198\dots$$
 (35)

(1018 exact digits are known [50]). Analogously to the Boltzmann–Gibbs case, a q-generalized Pesin identity $K_q = \lambda_q$ is verified. Consistently, K_q numerically appears to vanish for q above that special value and to diverge for q below that value. In other words, this is the unique value of the index q, for which $S_q(t)$ asymptotically increases *linearly* with time, thus yielding a *finite* entropy production per unit time.

• Still in the area of nonlinear dynamical systems, now with $N \ge 1$ degrees of freedom $\{x_1, x_2, ..., x_N\}$, scaling arguments [129] are consistent with

$$\frac{1}{1 - q_{\text{entropy}}} = \sum_{k=1}^{N} \frac{1}{1 - q_k},$$
(36)

where q_{entropy} is the value of q for which the entropy S_q production per unit time is *finite*, and $\{q_k\}$ characterize the q_k -exponential sensitivity to the initial conditions, respectively, associated with the variable x_k $(1 \ge q_1 \ge q_2 \ge \cdots \ge q_N > -\infty)$. This quite general relation is verified in the following illustrative cases:

(i) It suffices that one (q_1) of these *q*-indices equals unity and $q_{entropy} = 1$ hence $S_{q_{entropy}} = S_{BG}$;

(ii) If all $\{q_k\}$ are equal and smaller than unity, then:

$$q_{\text{entropy}} = 1 - \frac{1 - q_1}{N} \quad (q_1 < 1),$$
 (37)

hence $N \rightarrow \infty$ implies $q_{\text{entropy}} = 1$; Equations (30) and (31) as well as the model (with N exchangeable random variables) introduced and discussed in [106,130] correspond to this case;

(iii) If N = 1, then $q_{\text{entropy}} = q_1$, which corresponds to the cases (32)–(35) here above. Another important issue is the central limit theorem attractor in the space of distributions when time-averaging a single coordinate of the system. Within this theorem, $y \equiv \sum_{t=1}^{T} x_t$ is defined. For example, for the logistic map with a = 2 one has that, after proper scaling and centering, the distribution P(y) is given by a Gaussian, according to the classical central limit theorem, whereas, for $a = a_c$, it is seemingly given by a q-Gaussian with q > 1.

Let us further illustrate this issue with a paradigmatic *conservative* system, namely, the standard map, introduced by Chirikov in 1979. This area-preserving map turns out to be relevant within a variety of physical situations such as particle confinement in magnetic traps, particle dynamics in accelerators, comet dynamics, the ionization of Rydberg atoms, electron magnetotransport. It is defined as follows:

$$p_{i+1} = p_i - K \sin x_i \pmod{2\pi} \quad (K \ge 0)$$

$$x_{i+1} = x_i + p_{i+1} \pmod{2\pi}$$
(38)

Each (x, p) point yields a Lyapunov exponent $\lambda^{(1)} = -\lambda^{(2)} \ge 0$. Next, along the central limit theorem lines, let us define the following quantity:

$$\bar{y} \equiv \sum_{i=1}^{T} (x_i^{(j)} - \langle x \rangle) , \qquad (39)$$

with

$$\langle x \rangle \equiv \frac{1}{M} \frac{1}{T} \sum_{j=1}^{M} \sum_{i=1}^{T} x_i^{(j)}$$
, (40)

where $M \to \infty$ is the number of initial conditions and $T \to \infty$ is the number of iterations for each of those M initial conditions. The limiting K = 0 case (hence a linear map, though with a highly nontrivial set of stable orbits) exhibits zero Lyapunov exponent in the entire phase space. It has been studied [131] and the attractor is a *q*-Gaussian with

$$q = 2. (41)$$

This is in notorious variance with the attractor corresponding to $K \gg 1$, where the Lyapunov exponents corresponding to almost all the points of the phase space are positive, Boltzmann's molecular chaos hypothesis certainly is valid, and the corresponding attractor is a Gaussian, i.e., q = 1, in conformity with the standard central limit theorem.

• An interesting short-range-interacting *d*-dimensional ferromagnetic system is that whose symmetry is dictated by rotations in *n* dimensions, i.e., the so called O(n) symmetry (n = 2 corresponds to the *XY* model, n = 3 corresponds to the Heisenberg model, and so on; $n \to \infty$ corresponds to the spherical model). As soon as one focuses on the kinetics of point defects during a quenching from a high temperature to a zero temperature for the d = n model, its theoretical description can be done in terms of a time-dependent Ginzburg–Landau equation [132,133]. The distribution of the vortex velocity turns out to be a *q*-Gaussian with

$$q = \frac{d+4}{d+2}.\tag{42}$$

The index q decreases from two to one when d increases from zero to infinity. It is certainly intriguing, although yet unexplained, the fact that the value of q given in Equation (42) separates finite from infinite variance for d-dimensional q-Gaussians.

• For the coherent noise model for earthquakes and biological extinction, it has been possible to prove [134,135] the emergence of a *q*-Gaussian with

$$q = \frac{\tau + 2}{\tau}, \tag{43}$$

where τ is the index characterizing the (asymptotically) power-law distribution of the avalanches.

• Cold atoms. Lutz predicted in 2003 [136] that the velocity distributions of cold atoms in dissipative optical lattices should be *q*-Gaussians with

$$q = 1 + \frac{44 E_R}{U_0} \,, \tag{44}$$

where E_R and U_0 are, respectively, the recoil energy and the potential depth. The prediction was verified three years later [137,138].

Granular matter. The following scaling law was predicted in 1996 [84] (see also [83]):

$$\mu = \frac{2}{3-q},\tag{45}$$

where *q* is the index of the *q*-Gaussian distribution of d = 1 fluctuations and μ is the exponent associated with anomalous diffusion (i.e., the quadratic position x^2 scales like t^{μ} ; q = 1 recovers the known scaling for Brownian motion normal diffusion, i.e., $\mu = 1$). This scaling relation was experimentally verified in 2015 [139], within a $\pm 2\%$ precision along a wide experimental range, for vertically confined grains under horizontal shear.

• In some simple models, such as the so-called "inelastic Maxwell model," analytic calculations can be performed; see, e.g., [140]). The velocity distribution that is obtained, from the microscopic dynamics of the system of cooling experiments for a spatially uniform gas whose temperature is monotonically decreasing with time, is given by (see [141] and references therein) a *q*-Gaussian with

q

$$=3/2.$$
 (46)

Lattice Boltzmann models for fluids. The incompressible Navier–Stokes equation has been considered [121] on a discretized *D*-dimensional Bravais lattice of coordination number *b*. It is further assumed that there is a single value for the particle mass, and also for speed. The basic requirement for the lattice Boltzmann model is to be Galilean invariant (i.e., invariant under a change of inertial reference frame) like the Navier–Stokes equation itself. It has been proved [121] that an H-theorem is satisfied for a trace-form entropy only if it is *S_q* with

$$q = 1 - \frac{2}{D}$$
 (47)

Therefore, q < 1 in all cases (q > 0 if D > 2, and q < 0 if D < 2), and approaches unity from below in the $D \rightarrow \infty$ limit. This study has been generalized by allowing multiple masses and multiple speeds. Galilean invariance once again mandates [122] an entropy of the form S_q , with a unique value of q determined by a transcendental equation involving the dimension and symmetry properties of the Bravais lattice as well as the multiple values of the masses and of the speeds. Of course, Equation (47) is recovered for the particular case of a single mass and a single speed. Finally, Equation (47) may be seen as one more verification of the structure reflected in Equation (37).

• In the area of high-energy collisions, a highly interesting development based on firstprinciple Yang–Mills/QCD grounds became available a couple of years ago [142]. It yielded

$$\frac{1}{q-1} = \frac{11}{3}N_c - \frac{2}{3}N_f \,, \tag{48}$$

where N_c is the number of colors and N_f the number of flavors. If $N_c = N_f/2 = 3$, one finds 1/(q-1) = 7, hence $q = 8/7 \simeq 1.14$, which is amazingly close to the LHC (Large Hadron Collider) values indicated in the figure of [143] exhibiting the fittings (along very many experimental decades; notice also that the fitting temperature T = 0.13 GeV practically corresponds to the mass of pions, ubiquitous in proton-proton collisions) with which Kapusta begun his talk. Furthermore, if $N_c = N_f = 3$ is used, $q = 10/9 \simeq 1.11$ is obtained which, interestingly enough, coincides with the value phenomenologically advanced in 2000 by Walton and Rafelski [144] in their approach of a heavy quark diffusing in a quark–gluon plasma.

By the way, let me mention at this point that some years ago I briefly met in Brazil with the CERN researcher J. R. Ellis. I asked him whether he was aware that the q-exponential distribution ubiquitously emerged in high-energy collision data. He answered that he was aware. I then asked him whether he thought that this fact could be deduced from QCD. He answered that the value of q possibly yes, but the whole distribution probably not. The Deppman–Megias–Menezes relation (48) is by all means relevant along that line of thought; so are also the Wilk–Wlodarczyk, Beck and Beck–Cohen connections between q, fluctuations and degrees of freedom [145–147].

3. Final Remarks

Kapusta started his presentation by declaring "I am going to get to the bottom of this". He said that "*q* is not a fundamental or an intrinsic quantity to be interpreted in terms of new statistics" and "should not be considered a fundamental constant", and, while conceding that it constitutes an efficient fitting parameter, he concluded by stating that the

index *q* "is not a fundamental quantity *in any sense*". Along the same line, Giorgio Torrieri enthusiastically congratulated a "really great talk" and declared, in a very happy tone, "I will forward the video to many people".

Before ending, it might be useful to make explicit the following analogy. Within the Newtonian theory of the planetary system, there is *only one* fundamental fitting parameter, namely, the gravitational constant, *G*. All the rest about the orbits and motion of planets, asteroids, comets, etc., would be uniquely determined *if* (and what a big *if* is this one) we knew all the involved masses and their initial conditions at some time of the system's history. It is our human impossibility of having not only this information but also access to an unthinkably powerful computer to allow the corresponding Newtonian equations to be blamed if the task cannot be completed, certainly not the Newtonian theory. Still, this theory showed to be capable of reproducing the Kepler's empirical laws, in particular the elliptic analytic form of the orbits of the planets. Using these general mathematical forms, astronomers have been able to quite precisely determine the orbit of say Mars by introducing as many fitting parameters as necessary to close the description.

Similarly, nonextensive statistical mechanics has *only one* fundamental fitting parameter, namely, the Boltzmann constant, *k*, shared in fact with Boltzmann–Gibbs statistical mechanics. All the rest would be uniquely determined *if* (and what a big *if* is this one) we were able to analytically scrutinize the usually intractable mathematics associated with the microscopic (or at least the mesoscopic) dynamics of the nonlinear system so that the necessary indices *q* are determined from mechanical/electromagnetic first principles. It is our human impossibility of handling this formidable mathematics to be blamed if the task cannot be completed, certainly not the nonextensive thermostatistical theory! Still, this theory showed to be capable of providing useful analytical forms such as *q*-exponentials and *q*-Gaussians, as well as *H*-theorems, generalized central limit theorems, among others, which make possible the desired study of a specific complex system by introducing as many fitting parameters as necessary to close the description.

The present paper replies to Joseph Kapusta's talk and hopefully puts things in a further updated and more reliable perspective, proving in particular that the *q* indices surely are definitively more than possible convenient fitting parameters. As the reader has had the opportunity to directly verify here by themselves, the "bottom of this" is quite further deeper. Allow me to finish by reminding that, in fact, theoretical physics, like anything else, always appears to admit some new deeper steps down, below what was previously thought to be the bottom line (e.g., the nowadays celebrated emergence of relativity from Newtonian mechanics and, almost simultaneously, of quantum mechanics as well). Quoting Henry David Thoreau (*Walden; or, Life in the Woods*): "It is never too late to give up our prejudices. No way of thinking or doing, however ancient, can be trusted without proof. What everybody echoes or in silence passes by as true today may turn out to be falsehood tomorrow, mere smoke of opinion, which some had trusted for a cloud that would sprinkle fertilizing rain on their fields."

An oral reply [148] to the content of Kapusta's talk was delivered on April 12, 2022 at a seminar, hosted by the Santa Fe Institute, New Mexico, and chaired by Sidney Redner.

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